

HWRF physics packages

2016 Hurricane WRF Tutorial

Jan 2016 College Park, MD

Presented by Weiguo Wang

EMC/NCEP/NOAA

http://www.dtcenter.org/HurrWRF/users/docs/users_guide/HWRF_v3.7a_UG.pdf

Outline

- Overview of HWRF parameterizations
- Land surface model
- Surface layer physics
- Planetary Boundary Layer (PBL)
- Convection
- Microphysics
- Radiation
- Plan for 2016 operational HWRF upgrades

Overview

1. At the initial operational implementation in 2007, HWRF physics suite was closely following as GFDL hurricane model physics.

2. Roots of HWRF physics packages (3.7a)

NCEP GFS (PBL, convection, RRTMG)

NCEP mesoscale, ETA/NMMB (microphysics)

GFDL (surface physics)

WRF (LSM, cloud fraction)

...

Some modified for the tropical environment.

Thermodynamic equation

$$\frac{\partial T}{\partial t} = -u \frac{\partial T}{\partial x} - v \frac{\partial T}{\partial y} + \frac{P}{R} \omega \sigma + FT + \frac{\check{Q}}{C_P}$$

Dynamics
Physics

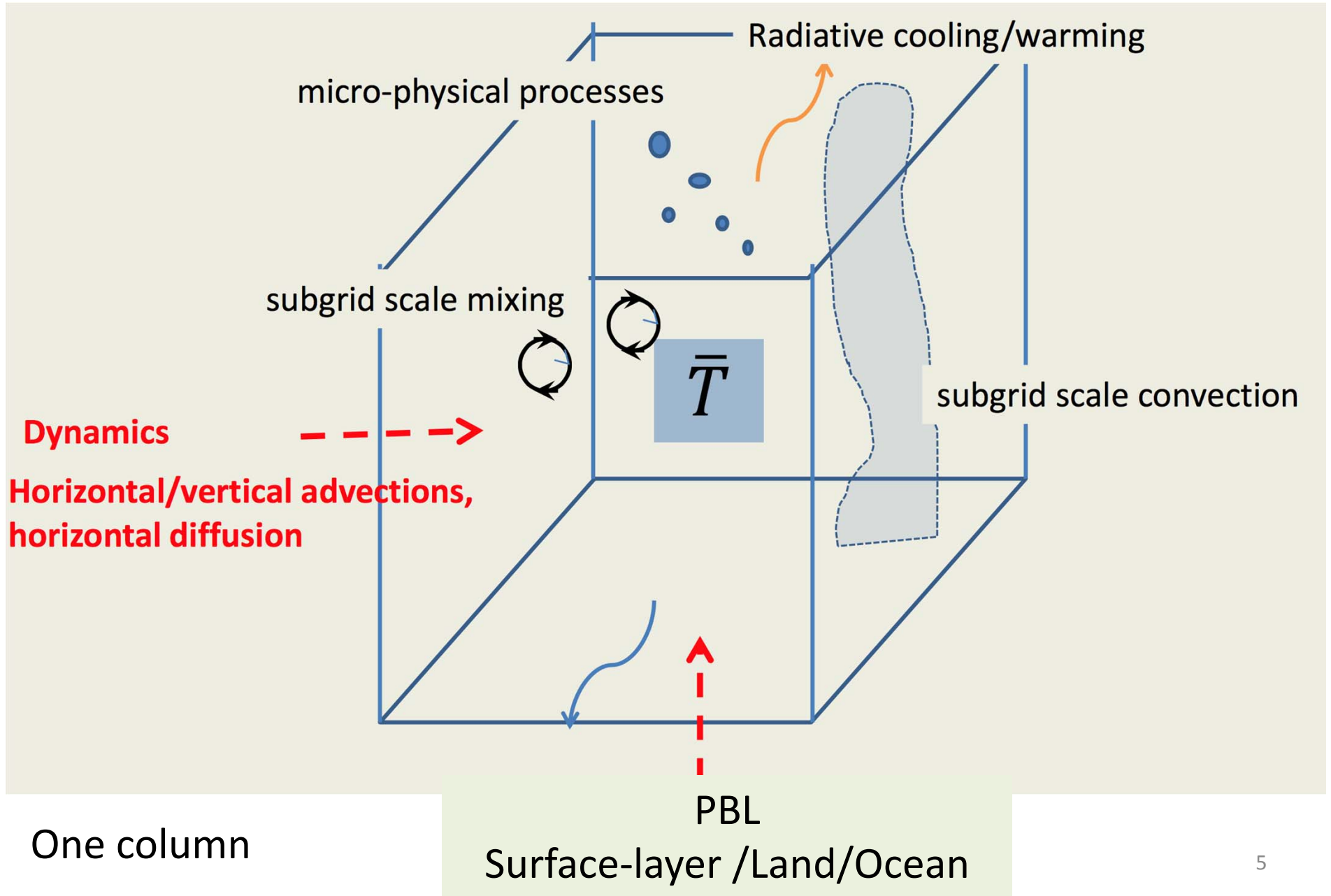
Time tendency
horizontal advection
vertical advec. + H. diffusion
adiabatic heating
diabatic heating

Tendency

where, $\sigma = - \left[\frac{RT}{P\theta} \right] \frac{\partial \theta}{\partial p}$

Diabatic heating: phase change of water
 Radiative processes
 Subgrid vertical mixing
 Surface fluxes

- convection, microphysics
- radiation
- PBL, convection
- air-sea interaction, land surface



Phys in 2015 operational HWRF

Scheme	Description
Land model	Community Noah land surface model (LSM)
Surface layer	M-O similarity, GFDL based code, cd, ch match obs.
PBL	GFS PBL + modification of K and Ric #
Convection	Simplified Arakawa-Schubert scheme with modifications #
Microphysics	Ferrier-Aligo scheme for high resolution model#
Radiation	RRTMG, partial cloudiness

#: planned upgrade for 2016

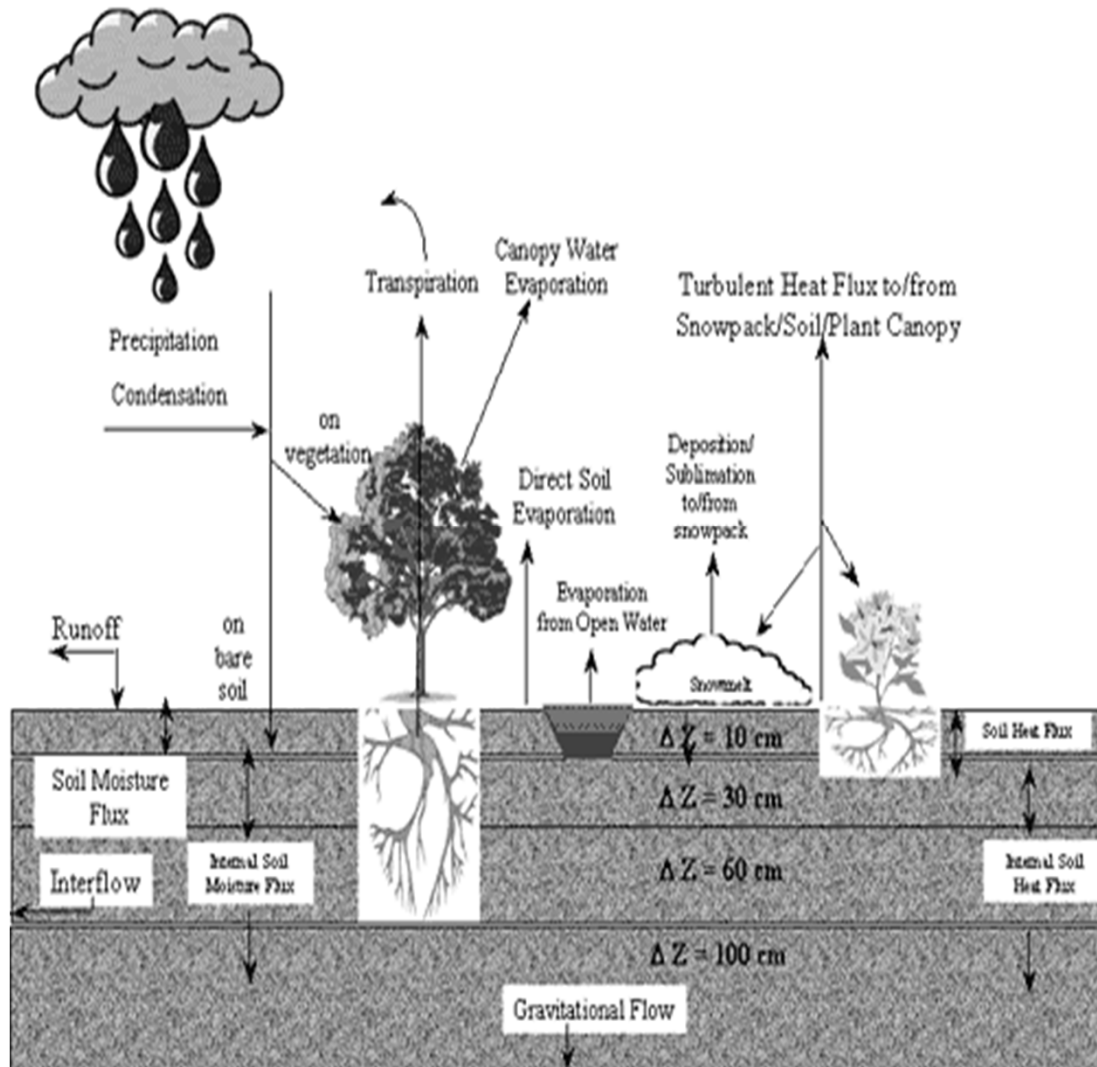
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Land surface model (LSM)

- **Why?** To provide heat and moisture fluxes over land points and sea-ice points. These serve as a lower boundary condition for the vertical transport done in the PBL schemes
- **How?** By using atmospheric information from the surface-layer scheme, radiative forcing from the radiation scheme, and precipitation forcing from the microphysics and convective schemes, in combination with layers of the soil and handle vegetation, root, and canopy effects and surface snow-cover prediction.

Noah LSM



<http://www.ral.ucar.edu/research/land/technology/lsm.php>

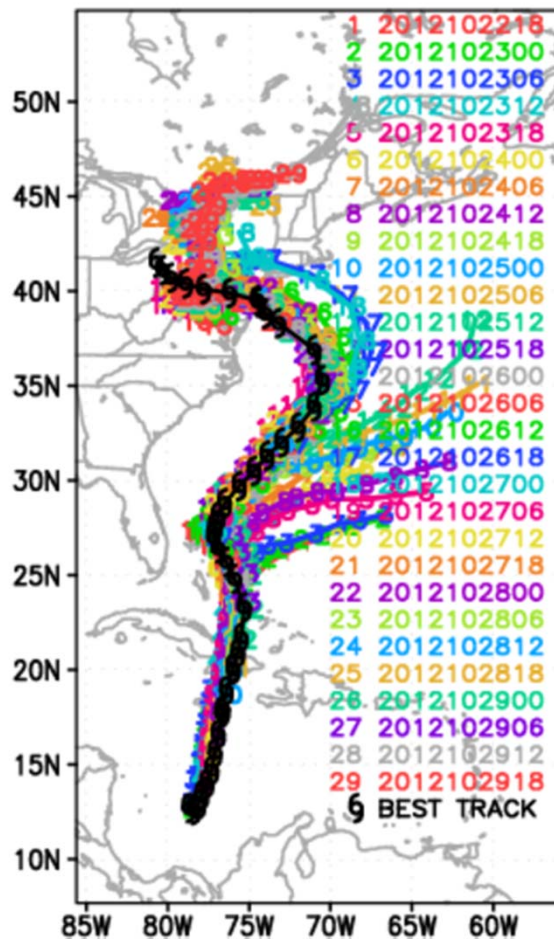
- NCEP/NCAR, Universities
- 1 canopy layer, 4-layer soil
- Prognostic variables:
 - Moisture and temperature in the soil layers
 - Water stored on the canopy
 - Snow stored on the ground.
- It includes root zone, evapotranspiration, soil drainage, and runoff, taking into account vegetation categories, monthly vegetation fraction, and soil texture.
- Many potential downstream applications with HWRF
- Used by HWRF since 2015

SANDY 18L 2012

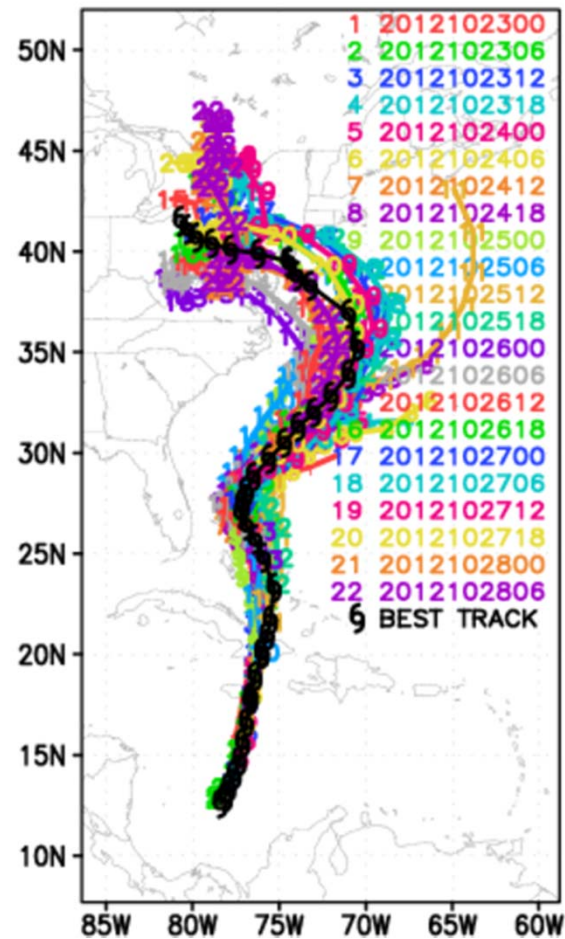
HWRF + slab land model

HWRF + NOAH LSM

H130 forecast: SANDY18L (al182012)



H133 forecast: SANDY18L (al182012)



Slab model: cold bias of TSK; NOAH LSM improved TSK & track simulations 10

LSM in Namelist

&physics

....

num_soil_layers = 4,
sf_surface_physics = 2, 2, 2,

....

Outline

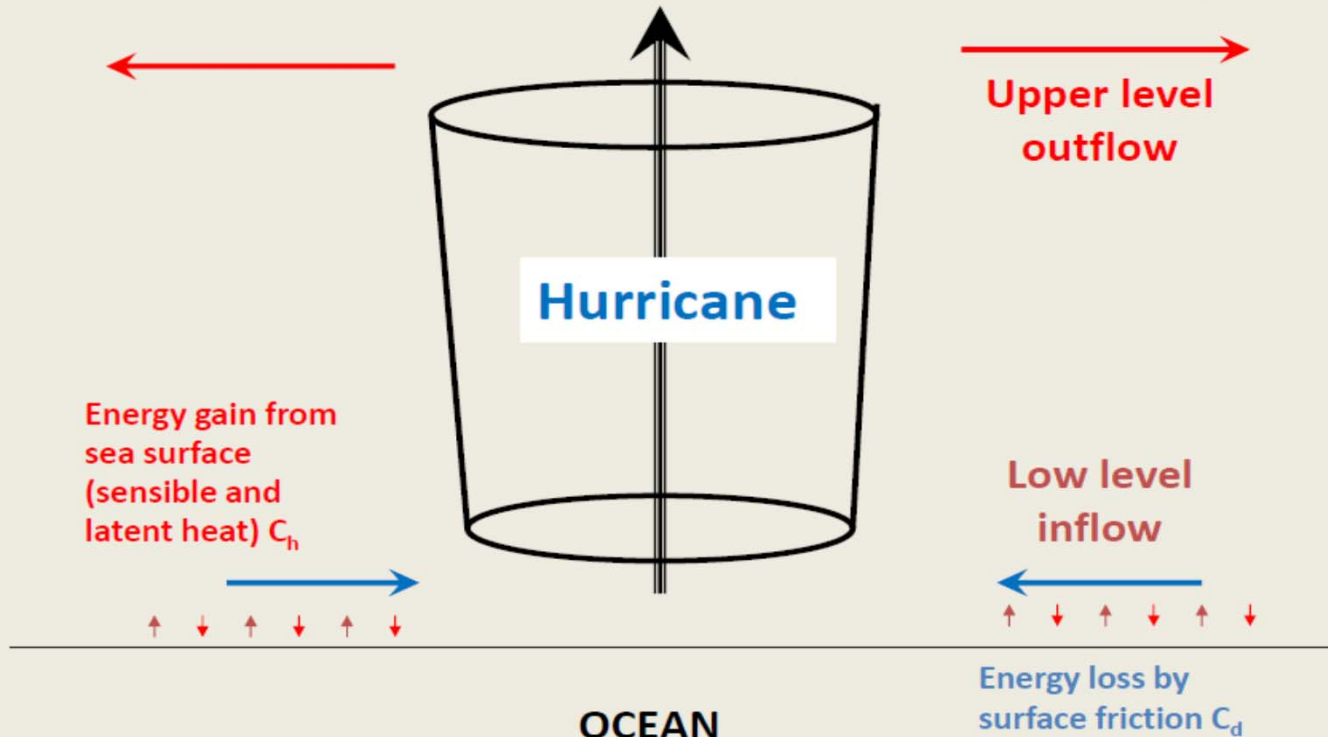
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Surface-layer physics

- **Why/what?** calculating stability functions, velocity scales (u_* , w_* ...), exchange coefficients, and surface fluxes
 - **How?** Monin-Obkuhov similarity theory
- Note:** provide no tendencies, only the stability-dependent information

Air-sea interactions

Hurricane intensity is proportional to $\sqrt{C_h/C_d}$ over ocean – Emanuel(1995)



Because main energy sources and sinks of tropical cyclones are sensible/latent heat fluxes over warm ocean and momentum flux (dissipation) over land, the determination of surface fluxes plays a critical role in predicting accurate hurricane intensity.

$$H_o = \rho C_p C_h |V| (\theta_s - \theta_a)$$

$$L_o = \rho L C_h |V| (Q_s - Q_a)$$

$$\vec{\tau}_o = \rho C_d |V| \vec{V}$$

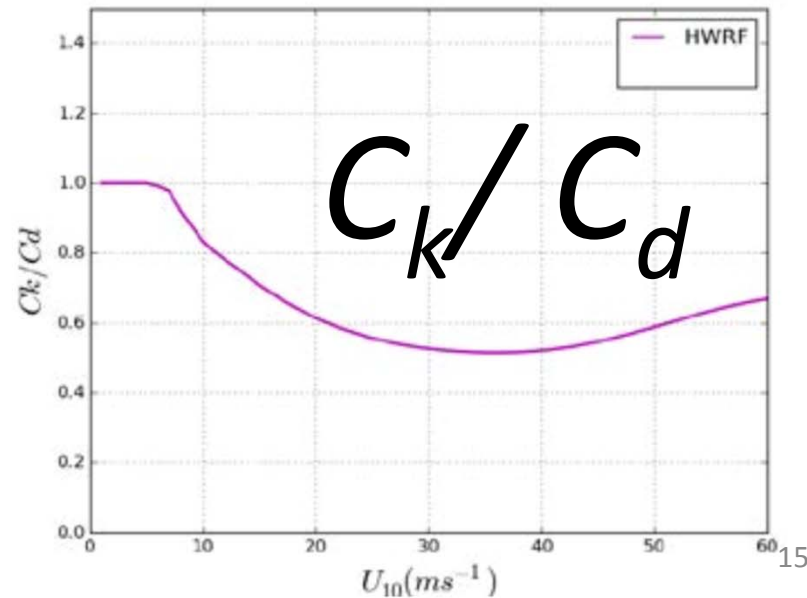
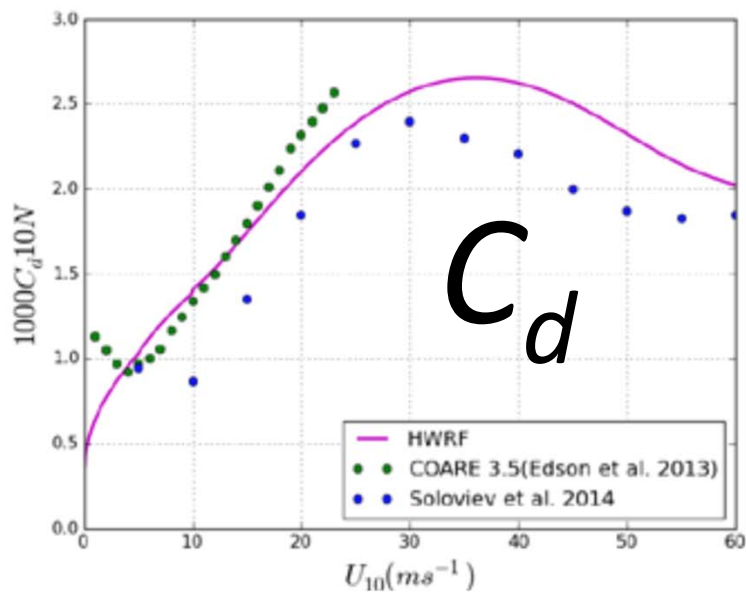
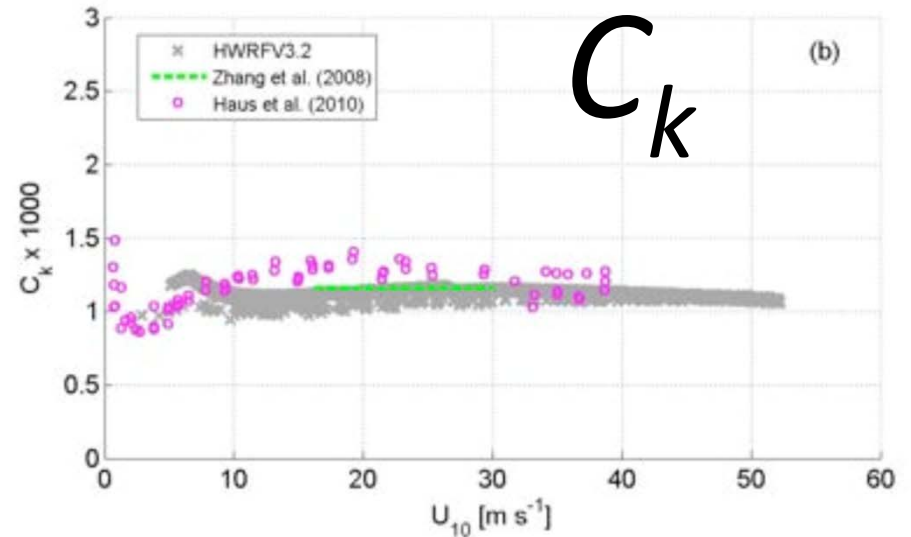
The air-sea flux calculations use a bulk parameterization based on the Monin-Obukhov similarity theory

Surface layer parameterization

$$C_d = \kappa^2 \left(\ln \frac{z_m}{z_0} \right)^{-2}$$

$$C_k = \kappa^2 \left(\ln \frac{z_m}{z_0} \right)^{-1} \left(\ln \frac{z_T}{z_0} \right)^{-1}$$

HWRF specifies z_0 as a function of wind to match obs C_d and C_k



Surface-layer in Namelist

&physics

....

`sf_sfclay_physics = 88, 88, 88,` ← GFDL sfc

`icoef_sf = 2, 2, 2,` ← Choose different Cd, Ck

`lcurr_sf = F, F, F,` ← For ocean, future

....

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Planetary boundary layer (PBL) parameterization

- **Why?** To represent the mixing effects of eddies too small to be solved by the dynamical core
- **How?** It determines flux profiles and provides atmospheric tendencies (in the entire atmospheric column) of: temperature, moisture (including clouds)
Horizontal momentum

$$\frac{\partial \Phi}{\partial t} = - \frac{\overline{\partial w' \phi'}}{\partial z}$$

$$\overline{w' \phi'} = -K \frac{\partial \bar{\phi}}{\partial z} + K \gamma \quad \leftarrow \text{Counter-gradient term}$$

$$\overline{w' \phi'} = \text{local flux} + \text{non local flux}$$

Turbulent eddy diffusivity , K

$$K = l^2 f(Ri) \left| \frac{\partial \bar{u}}{\partial z} \right|$$

Stable conditions,
or above PBL top

$$K = \kappa \left(\frac{u_*}{\Phi_m} \right) \left[\alpha \left(1 - \frac{z}{h} \right)^2 \right]$$

Unstable conditions
below PBL top

α : K adjustment coef

$$h = Ri_c \frac{\theta_{VS} U^2(h)}{g(\theta_V(h) - \theta_S)}$$

PBL height depends on Ri_c , wind, θ_V ,
and θ_S at ground and PBL top

$$Ri_c = 0.16 (10^{-7} R_o)$$

Critical Richardson number a function
of the local Rossby Number

$$R_o = \left(\frac{U_{10}}{fz_o} \right)$$

Rossby number depends on 10-m
speed and roughness length

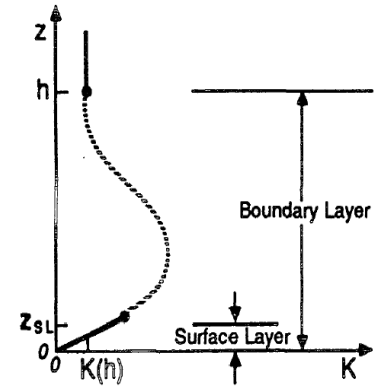


FIG. 1. Typical variation of eddy viscosity K with height in the boundary layer proposed by O'Brien (1970). Adopted from Stull (1988).

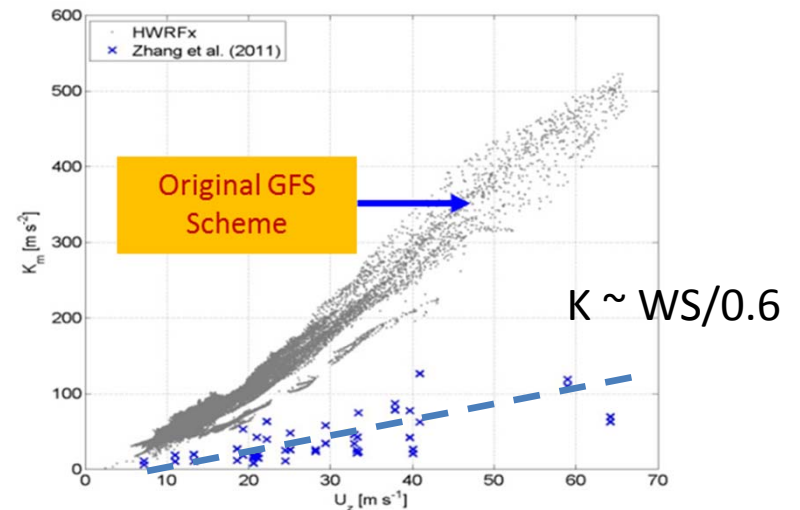
nd Pan (1996)

K adjustment

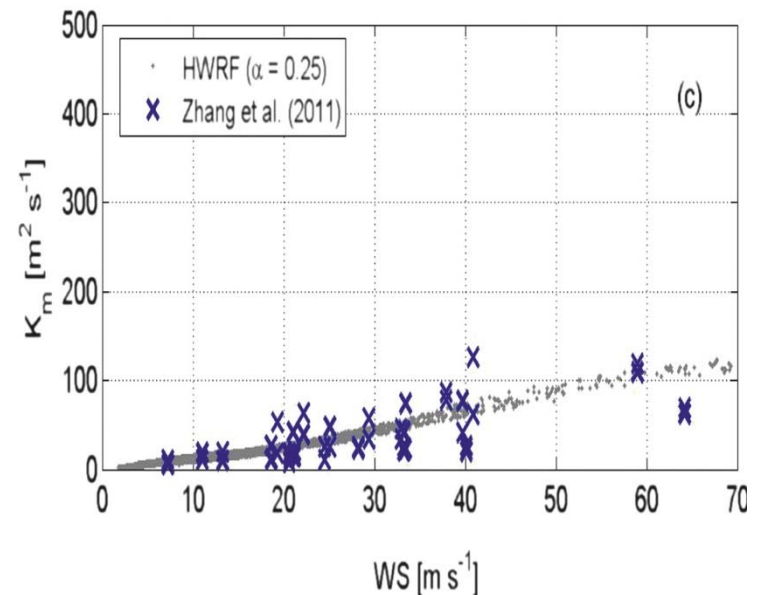
Issue: Eddy diffusivity from regular GFS PBL scheme was too large.

Solutions:

- Upgrade for 2013 operational HWRF: introduce a variable (α) to modulate K (Gopal et al, 2012, 2013)
- Upgrade for 2014 operational HWRF: make critical Richardson number dependent on Rossby number when computing PBL height
- Upgrade for 2015 operational HWRF: set upper boundary for K at $\sim 500\text{m}$ to $\text{WindSpeed}/0.6$ (Fovell et al, 2015)



Gopal et al.
2012;2013 20



PBL in Namelist

&physics

....

bl_pbl_physics = 3, 3, 3,

var_ric = 1.0,

coef_ric_l = 0.16,

coef_ric_s = 0.25,

gfs_alpha = -1.0, -1.0, -1.0,

....

1: Use variable Ric# for h; 0:no

Coefficient of Ric# : land point

Coefficient of Ric# : sea point

K adjustment.

alpha=abs(gfs_alpha)

<0, Use ws/0.6 method

>0, Use a constant alpha

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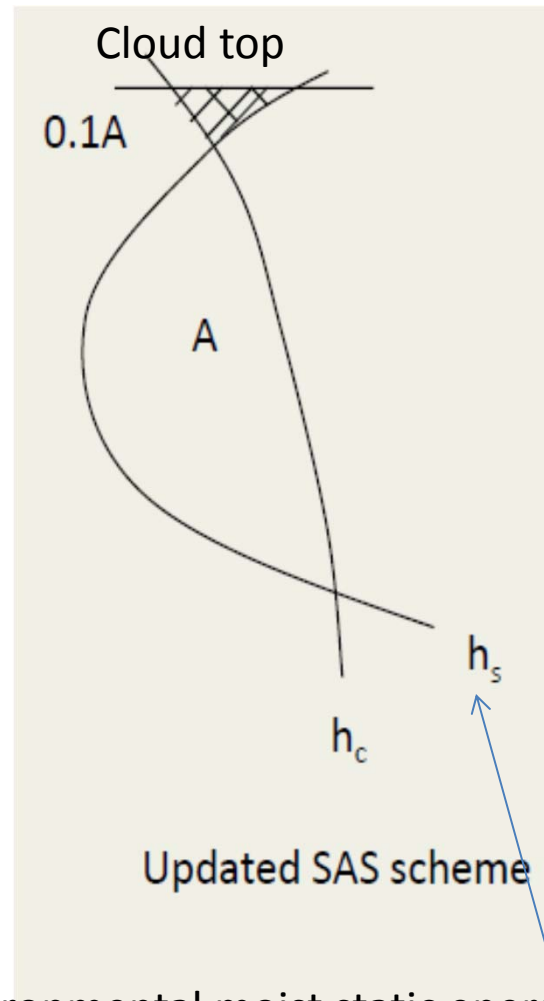
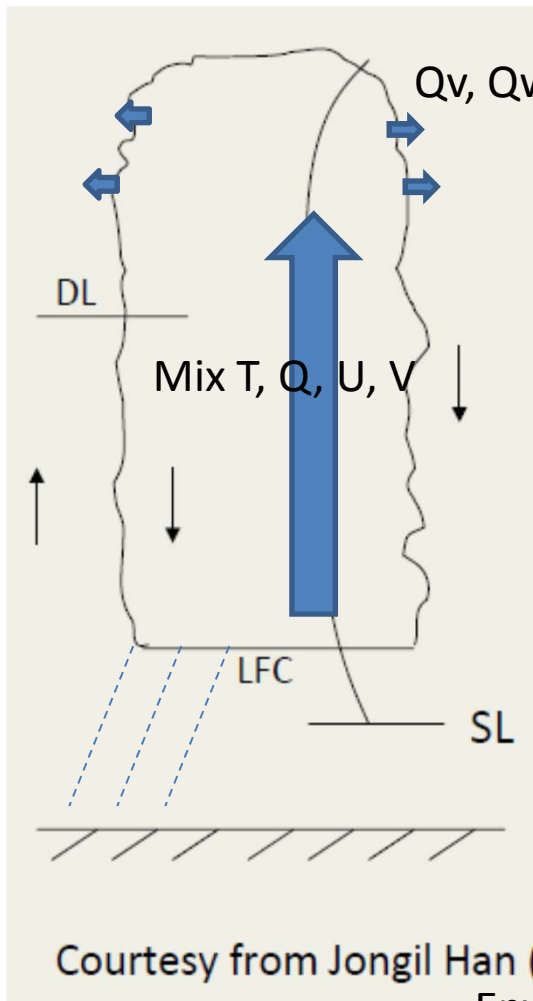
Convective parameterization

(also referred to as cumulus parameterization scheme)

- **Why?** To represent the convection too small to be resolved explicitly given the grid spacing. Convection has the effect of reducing the thermodynamic instability
- **How?** Mass flux scheme
 - Vertically transporting:
 - Heat
 - Momentum
 - Moisture

Simplified AS (SAS)

- Arakawa and Schubert (1974) and simplified by Grell (1993)
- Also used in the GFS and GFDL models



Features

- Cloud size \ll grid size
- One cloud top
- Deep: thickness > 150 hpa; shallow: < 150
- Triggered when cloud-work function (\sim CAPE) $>$ threshold
- T, Q adjusted, mass flux method
- Simple microphysics, evaporation, detrain condensation/vapor
- Momentum mixing

CU in Namelist

&physics

....

cu_physics = 84, 84, 0,
mommix = 1.0, 1.0, 1.0,
ncnvc = 2, 6, 6,
sas_pgcon = 0.55, 0.2, 0.2,
sas_mass_flux = 0.5, 0.5, 0.5,

....

No convection used in D03

Fraction of momentum mixing

Calling frequency, time steps

Convectively forced pressure gradient.

Mass flux limit

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Microphysics

- **Why?** To represent grid-scale microphysical processes such as
 - Aggregation, accretion, growth, precipitation, etc
 - Phase change and latent heat release due to condensation, evaporation, deposition, sublimation, freezing, melting.
- **How?** Predicting one or more moments of the distribution of various microphysics species such as rain, cloud, ice

HWRF uses a single moment scheme

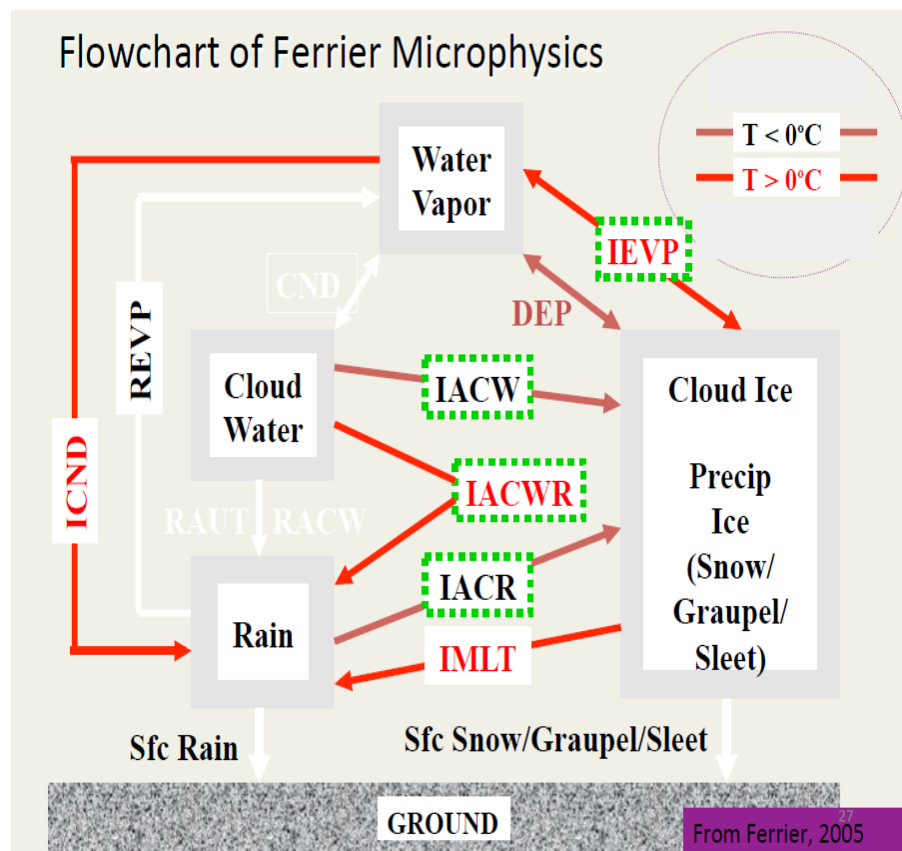
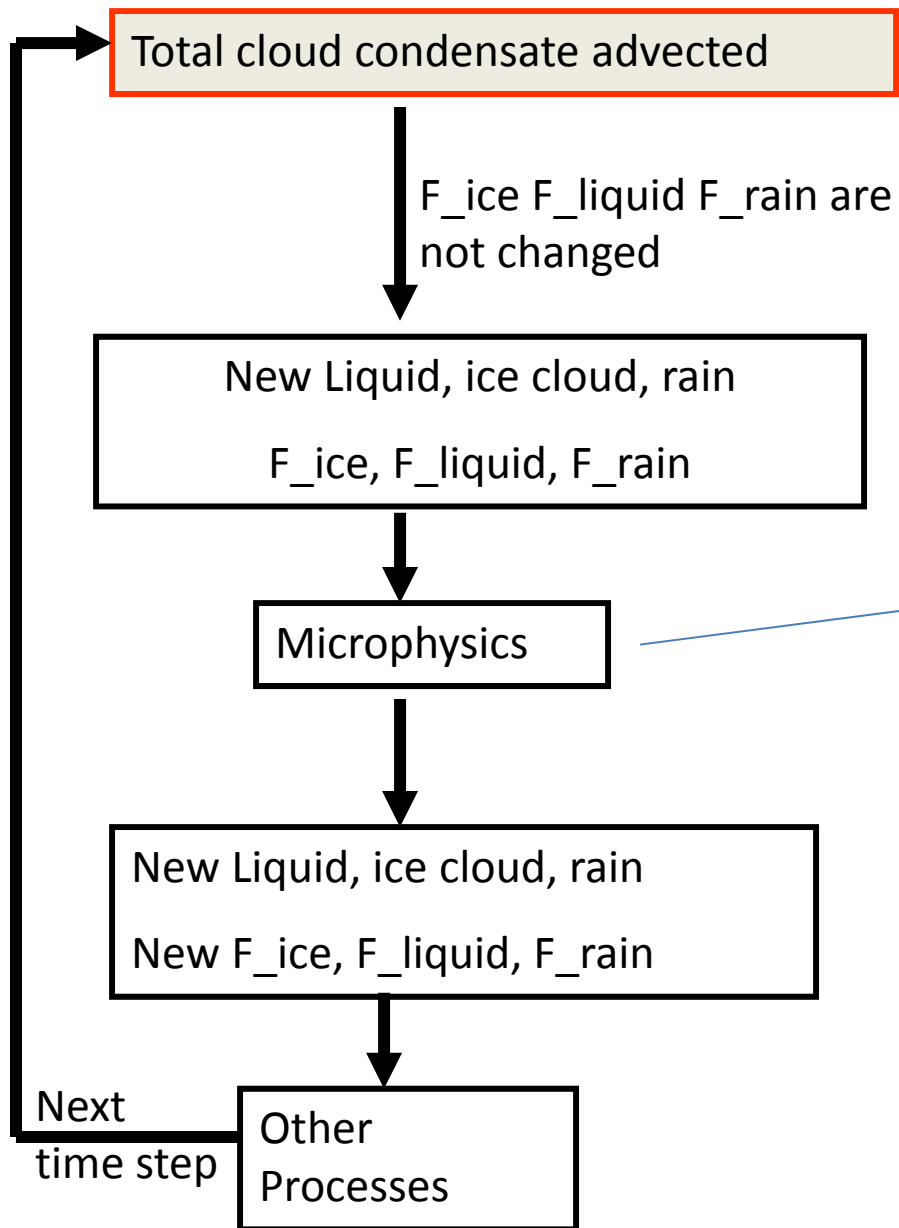
Ferrier-Aligo scheme (NMMB,HWRF) ,

modified from Ferrier/eta scheme

Ferrier-Aligo scheme

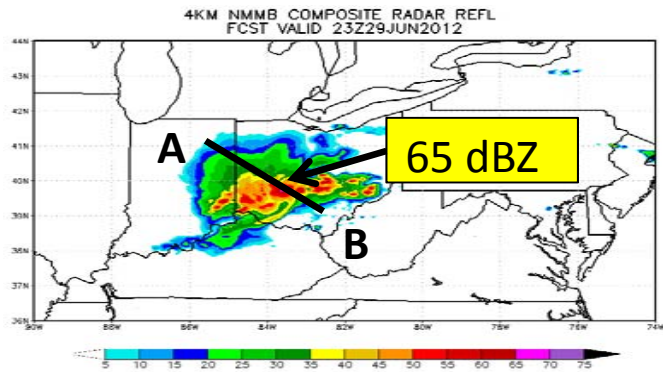
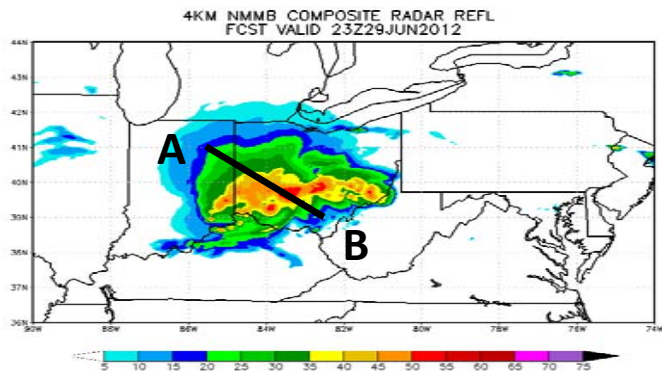
- Designed for efficiency
- Predicts mixing ratios; Diagnoses number concentrations
- Advection only of total condensate (CWM) and vapor
- Diagnostic cloud water, rain, & ice (ice/snow/graupel)
- Fractions of cloud water & ice within a column are fixed during advection
- Variable density for precipitation ice
- (snow/graupel/sleet) – “rime factor” (F_{rime})
- Maximum number concentration of large ice varies in different cloud regimes. Promote supercooled liquid water.

Flow chart

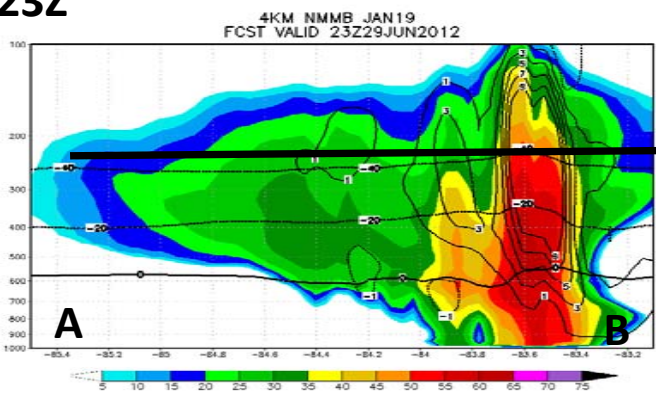
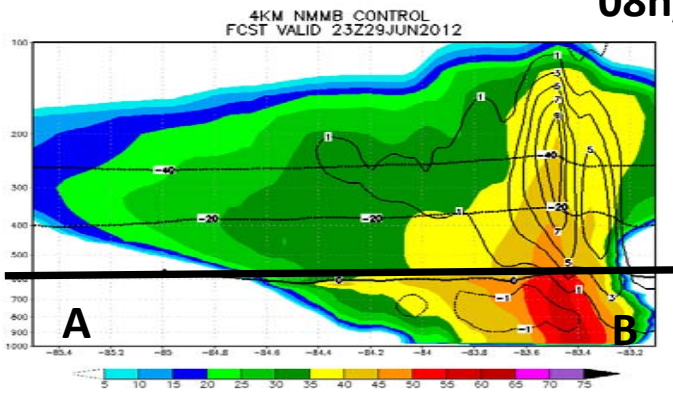


Ferrier/eta Microphysics

Ferrier-Aligo Microphysics



08h/23Z



50 dBZ
reflectivity below
600 mb

50+ dBZ
reflectivity up
to
250mb

Courtesy Eric Aligo

MP in Namelist

&physics

....

mp_physics = 5, 5, 5,

nphs = 2, 6, 6,

....

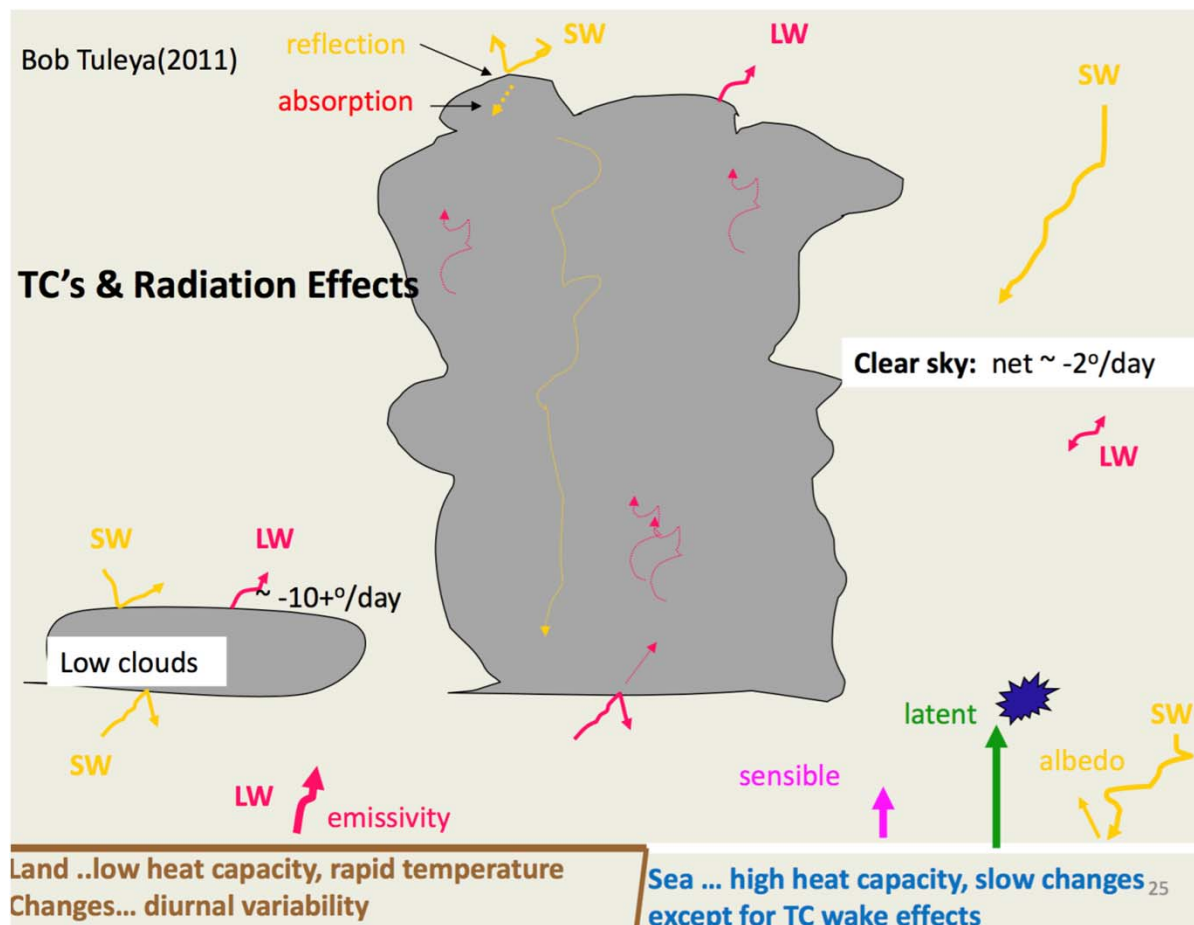
Call frequency in time steps

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Radiation

- **Why?** To represent the heating and cooling due to short and long wave radiation



Rapid Radiative Transfer Model (RRTMG)

- Modified from RRTM (Iacono et al. 2008)
- Long wave includes absorptions of:
 - water vapor, carbon dioxide, ozone, methane, nitrous oxide, oxygen, nitrogen, and the halocarbons
- Short wave includes absorptions of:
 - water vapor, carbon dioxide, ozone and methane
- Calculations are made over spectral bands, with:
 - longwave (16 bands)
 - shortwave (14 bands)
- Uses the Monte Carlo Independent Cloud Approximation random Overlap method to overlap clouds
- Uses a cloud fraction scheme developed by Thompson (NCAR) that represents sub-grid scale clouds
- Ice particle size depends on temperature, instead of constant (NMMB tests).

Radiation in Namelist

&physics

....

ra_lw_physics = 4, 4, 4,

ra_sw_physics = 4, 4, 4,

icloud = 3,

nrads = 30, 90, 270,

nradl = 30, 90, 270,

....

Cloud fraction in Radiation

Call frequency of SW module

Call frequency of LW module

Full Physics Options in 2015 operational HWRF

Can be configured through “hwrf.conf”

&physics

num_soil_layers = 4,
mp_physics = 5, 5, 5,
ra_lw_physics = 4, 4, 4,
ra_sw_physics = 4, 4, 4,
sf_sfclay_physics = 88, 88, 88,
sf_surface_physics = 2, 2, 2,
bl_pbl_physics = 3, 3, 3,
cu_physics = 84, 84, 0,
mommix = 1.0, 1.0, 1.0,
var_ric = 1.0,
coef_ric_l = 0.16,
coef_ric_s = 0.25,
h_diff = 1.0, 1.0, 1.0,
gwd_opt = 2, 0, 0,
sfenth = 0.0, 0.0, 0.0,
nrads = 30, 90, 270,
nradl = 30, 90, 270,
nphs = 2, 6, 6,
ncnvc = 2, 6, 6,
ntrack = 9, 9, 18,

gfs_alpha = -1.0, -1.0, -1.0,
sas_pgcon = 0.55, 0.2, 0.2,
sas_mass_flux = 0.5, 0.5, 0.5,
co2tf = 1,
vortex_tracker = 2, 2, 7,
nomove_freq = 0, 6, 6,
tg_option = 1,
ntornado = 2, 6, 18,
ens_cdamp = 0.2,
ens_pblamp = 0.2,
ens_random_seed = 99,
ens_sasamp = 50.0,
icloud = 3,
icoef_sf = 2, 2, 2,
lcurr_sf = F, F, F,
pert_cd = F,
pert_pbl = F,
pert_sas = F,

Other options

Limited tested:

- Convection schemes: Tiedtke, NSAS, KF, etc, tested by DTC
- Radiation: GFDL
- PBL: MYJ
- LSM: slab
- Surface layer: eta similarity
- Microphysics: Thompson, WSM6, Ferrier

Plan for 2016 upgrades

- GFS PBL EDMF scheme
- Scale-aware /Aerosol-aware SAS scheme
- Advection of Ferrier MP species.
- Thompson MP being tested.

GFS PBL EDMF

- Latest version GFS PBL

$$\overline{w' \phi'} = \text{local flux} + \text{non local flux}$$

HWRF 2015

Counter-gradient method to represent nonlocal flux

$$\overline{w' \phi'} = -K \left(\frac{\partial \bar{\phi}}{\partial z} - \gamma \right)$$

TKE based
dissipation heating

HWRF 2016

$$\overline{w' \phi'} = -K \frac{\partial \bar{\phi}}{\partial z} + M (\phi_w - \bar{\phi})$$

Local Eddy

Mass Flux

Scale-aware SAS

- For cumulus convection, a scale-aware parameterization will be necessary for the grid sizes of 500m ~ 10 km where the strong updrafts are partially resolved.

$$m'_b = (1 - \sigma_u)^2 m_b$$

Scale function

Based on Arakawa & Wu (2013): $\overline{w' \psi'} = (1 - \sigma_u)^2 \overline{(w' \psi')}_E$

σ_u : updraft area fraction (0~1.0)

m_b : original cloud base mass flux from AS quasi-equilibrium closure

m'_b : updated cloud base mass flux with a finite σ_u

Thanks!